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Numerical Relativity in higher dimensional spacetimes – Part 2

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Recap – what have we learned last time

- formulation of Einstein's equations

$${}^{(D)}R_{MN} - \frac{1}{2}g_{MN}{}^{(D)}R = 8\pi G_D T_{MN}$$

as time evolution problem

- $(D-1) + 1$ split of spacetime (\mathcal{M}, g_{MN})

$$\mathcal{M} = \Sigma_t \times \mathbb{R}$$

- line element

$$ds^2 = -(\alpha^2 - \beta_{\bar{m}}\beta^{\bar{m}}) dt^2 + 2\gamma_{\bar{m}\bar{n}}\beta^{\bar{m}} dt dx^{\bar{n}} + \gamma_{\bar{m}\bar{n}} dx^{\bar{m}} dx^{\bar{n}}$$

with

- coordinate degrees of freedom: α – lapse, $\beta^{\bar{m}}$ – shift vector
- dynamical variables:
 $\gamma_{\bar{m}\bar{n}}$ – (D-1)-metric, $K_{\bar{m}\bar{n}} = -\frac{1}{2}\mathcal{L}_n\gamma_{\bar{m}\bar{n}}$ – extrinsic curvature

Recap – what have we learned last time

$(D - 1) + 1$ split of Einstein's equations

- constraint subsystem

$$\mathcal{H} = {}^{(D-1)}R - K_{\bar{m}\bar{n}}K^{\bar{m}\bar{n}} + K^2 - 16\pi G_D\rho = 0$$

$$\mathcal{M}_{\bar{m}} = D_{\bar{n}}K^{\bar{n}}_{\bar{m}} - D_{\bar{m}}K - 8\pi G_D j_{\bar{m}} = 0$$

- evolution subsystem

$$\partial_t \gamma_{\bar{m}\bar{n}} = -2\alpha K_{\bar{m}\bar{n}} + \mathcal{L}_\beta \gamma_{\bar{m}\bar{n}}$$

$$\begin{aligned} \partial_t K_{\bar{m}\bar{n}} = & -D_{\bar{m}}D_{\bar{n}}\alpha + \alpha \left({}^{(D-1)}R_{\bar{m}\bar{n}} - 2K^{\bar{k}}_{\bar{m}}K_{\bar{k}\bar{n}} + KK_{\bar{m}\bar{n}} \right) \\ & + 4\pi G_D \alpha (\gamma_{\bar{m}\bar{n}}(S - \rho) - 2S_{\bar{m}\bar{n}}) + \mathcal{L}_\beta K_{\bar{m}\bar{n}} \end{aligned}$$

- change of variables and constraint addition
 \Rightarrow numerically stable evolution scheme, e.g., BSSN
- gauge choice – moving punctures

$$\partial_t \alpha = -2\eta_\alpha \alpha K$$

$$\partial_t \beta^{\bar{m}} = \frac{D-1}{2(D-2)} \zeta_\Gamma \tilde{\Gamma}^{\bar{m}} - \eta_\beta \beta^{\bar{m}} + \beta^{\bar{k}} \partial_{\bar{k}} \beta^{\bar{m}}$$

- $(D - 1) + 1$ formulation numerically too expensive for $D > 4$
- number of gridpoints $\sim N^{D-1}$
- code dealing with generic spacetime dimension D
vs. individual code for each dimension
- consider symmetric problems
→ reduction to $(3 + 1)$ or $(2 + 1)$ problems

1 Approach I – The Cartoon method

2 Approach II – Dimensional reduction by isometry

The Cartoon method



Cartoon method in $D = 4$

- goal: evolution of highly symmetric problems, e.g., with axis-symmetry (2+1)
 - natural choice: cylindrical coordinates
 - coordinate singularity at axis
 - sometimes tricky to handle
 - Cartesian coordinates
 - avoid coordinate singularities
 - requires 3+1 \rightarrow more expensive: N^3 instead of N^2 grid points
- \Rightarrow combine advantages

\Rightarrow **Cartoon method** (“CARTesian TWOdimensional”) (Alcubierre et al '99)

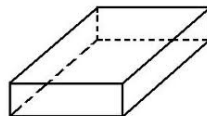
Cartoon method in $D = 4$

- for now: consider axissymmetric problem in $x - z$ -plane ($y = 0$), e.g., head-on collision along z -axis
- in $3d$: symmetry not explicit
 - spatial derivatives in all directions; in general: $\partial_y u \neq 0$
 - e.g., 4th order finite difference stencils

$$\partial_y u_j = \frac{1}{12dy} (u_{j-2} - 8u_{j-1} + 8u_{j+1} - u_{j+2})$$

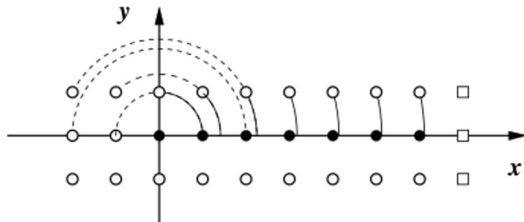
\Rightarrow require $2g + 1 = 5$ points in y -direction

- reduce $3d$ cube with N^3 grid points to $3d$ slab with $N^2 \cdot (2g + 1)$ grid points



Cartoon method in $D = 4$

- in interior of slab \rightarrow centered finite difference stencils
- at outer boundary \rightarrow physical boundary conditions
- evaluate functions at grid points $(x, y = 0, z)$
- $U(1)$ symmetry around z
 \Rightarrow obtain values at $y \neq 0$ by rotation and interpolation from field at $y = 0$
- required point $(\rho, 0, z) = (\sqrt{x^2 + y^2}, 0, z)$ might not be a grid point
 \Rightarrow interpolation



(Alcubierre et al '99)

Cartoon method in $D = 4$

- cylindrical coordinates (ρ, ϕ, z) with

$$x = \rho \cos \phi, \quad y = \rho \sin \phi, \quad z = z, \quad \rho = \sqrt{x^2 + y^2}$$

- consider rotation by angle ϕ_0

$$\rho' = \rho, \quad \phi' = \phi + \phi_0, \quad z' = z$$

- define diffeomorphism

$$R : ({}^3\Sigma_t \rightarrow ({}^3\Sigma_t \text{ with } R : T \rightarrow R^*T, \text{ for } T \in ({}^3\Sigma_t$$

$$(R(\phi_0)^i_j) = \begin{pmatrix} \cos \phi_0 & -\sin \phi_0 & 0 \\ \sin \phi_0 & \cos \phi_0 & 0 \\ 0 & 0 & 1 \end{pmatrix} = \begin{pmatrix} \frac{x}{\rho} & -\frac{y}{\rho} & 0 \\ \frac{y}{\rho} & \frac{x}{\rho} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

- R is symmetry if $R^*T = T$

Cartoon method in $D = 4$

- transformation of tensor $T(p) \rightarrow R^* T(R(p))$ (with $R^* T = T$)

$$T^{i_1 \dots i_p}_{j_1 \dots j_q}(x, y, z) = R^{i_1}_{k_1} \dots R^{i_p}_{k_p} (R^{-1})^{j_1}_{l_1} \dots (R^{-1})^{j_q}_{l_q} T^{k_1 \dots k_p}_{l_1 \dots l_q}(\rho, 0, z)$$

\Rightarrow evaluate functions @ $p' = R(p) = (x, y, z)$ from functions @ $p = (\rho, 0, z)$

- examples (\diamond):

- scalar

$$\Psi(x, y, z) = \Psi(\rho, 0, z)$$

- vector $V^i(x, y, z) = R^i_j V^j(\rho, 0, z)$

$$V^x(x, y, z) = \frac{x}{\rho} V^x(\rho, 0, z) - \frac{y}{\rho} V^y(\rho, 0, z)$$

$$V^y(x, y, z) = \frac{y}{\rho} V^x(\rho, 0, z) + \frac{x}{\rho} V^y(\rho, 0, z)$$

$$V^z(x, y, z) = V^z(\rho, 0, z)$$

- tensor

$$S_{ij}(x, y, z) = R^k_i R^l_j S_{kl}(\rho, 0, z)$$

Cartoon method in $D \geq 4$

- so far:
considered 3d space with $U(1)$ symmetry, e.g., head-on collision along z-axis
- further examples in $D > 4$:
(Yoshino & Shibata '09, Shibata & Yoshino '10,
Lehner & Pretorius '10, Okawa et al '11)
 - $D = 5$ with $U(1)$ symmetry, e.g., BH collisions with impact parameter

Cartoon method in $D = 5$

$D = 5$ with $U(1)$ symmetry, e.g., BH collisions with impact parameter

- consider Cartesian coordinates (x, y, z, w) with $(x, y, z = w)$
 - $x - y$ -plane as symmetry plane
 - Killing vector: $\partial_\psi = z\partial_w - w\partial_z$
- **Cartoon method**
 - simulate $D = 5$ problem in $3 + 1$
 $\Rightarrow (2g + 1)N^3$ grid points instead of N^4
 - consider cylindrical coordinates

$$x = x, \quad y = y, \quad z = \rho \cos \psi, \quad w = \rho \sin \psi$$

Cartoon method in $D = 5$

$D = 5$ with $U(1)$ symmetry (continued)

- strategy:

- 1 prepare data @ $(x, y, z, w = 0)$
 - 2 generate data @ $(x, y, z, w \neq 0)$ by rotation of data @ $(x, y, \rho, 0)$
- Killing vector: $\partial_\psi = z\partial_w - w\partial_z$
 - rotation by $\psi = \psi_0$ with rotation matrix:

$$(R(\psi_0)^{\bar{m}}_{\bar{n}}) = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & \cos \psi_0 & -\sin \psi_0 \\ 0 & 0 & \sin \psi_0 & \cos \psi_0 \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & \frac{z}{\rho} & -\frac{w}{\rho} \\ 0 & 0 & \frac{w}{\rho} & \frac{z}{\rho} \end{pmatrix}$$

- construct all necessary derivatives / quantities
 \Rightarrow happy number crunching...

Cartoon method

- so far:
considered 3d space with $U(1)$ symmetry, e.g., head-on collision along z-axis
- further examples in $D > 4$:
(Yoshino & Shibata '09, Shibata & Yoshino '10,
Lehner & Pretorius '10, Okawa et al '11)
 - $D = 5$ with $U(1)$ symmetry, e.g., BH collisions with impact parameter
 - $D = 5$ with $U(1) \times U(1)$ symmetry, e.g., Myers-Perry BH with 2 spin parameters (◇)
 - $D = 5$ with $SO(3)$ symmetry, e.g., head-on collision along z-axis (◇)
 - $D \geq 5$ spacetime with $SO(D-3)$ symmetry (“modified Cartoon method”) (◇)
- note: see lecture notes for examples

Modified Cartoon method in $D \geq 5$

consider D dimensional vacuum spacetime with $SO(D-3)$ symmetry (Shibata & Yoshino '10)

- coordinates $(t, x, y, z, w_1, \dots, w_{D-4})$

- line element

$$ds^2 = -(\alpha^2 - \beta^k \beta_k) dt^2 + 2\gamma_{kl} \beta^k dt dx^l + \gamma_{ab} dx^a dx^b + \gamma_{nn} d\Omega_{D-4}^2$$

- $x^a = (x, y, \rho)$ with $\rho = \sqrt{z^2 + \sum_{q=1}^{D-4} w_q^2}$

\Rightarrow symmetry around z, w_1, \dots, w_{D-4}

\Rightarrow geometric quantities depend only on t, x, y, ρ

- strategy:

- 1 solve for data in (x, y, z) hyperplane, with $w_1 = \dots = w_{D-4} = 0$
- 2 rotate data around hyperplane \Rightarrow data at $(x, y, z, w_1 \neq 0, \dots, w_{D-4} \neq 0)$

Modified Cartoon method in $D \geq 5$

consider D dimensional vacuum spacetime with $SO(D-3)$ symmetry (continued)

- straightforward Cartoon method
 - becomes numerically more expensive with increasing dimension
 - $N^3(2g+1)^{D-4}$ grid points
 - instead: make use of $SO(D-3)$ symmetry
 - \Rightarrow relations for tensor Q_{MN} , vector Q^M , scalar Q quantities (\diamond):
 - $Q_{ww} := Q_{w_1 w_1} = \dots = Q_{w_{D-4} w_{D-4}}$
 - $Q^{wq} = 0$, $Q_{aw_q} = 0$, $Q_{w_q w_r} = 0$ for $(r \neq q)$
 -
 - \Rightarrow all quantities and derivatives in $D-1$ dimensions are replaced by quantities and derivatives in the (x, y, z, w) plane
- now: happy number crunching ...

Dimensional Reduction by isometry

Dimensional reduction

dimensional reduction well known in physics

- Kaluza - Klein theory (1921)
 - ⇒ unification of gravity and gauge theories
 - e.g.: GR – Maxwell in $D = 4 \leftrightarrow$ GR in $D = 5$
- string theory compactifications
- Here:
 - “Killing reduction” – dimensional reduction by isometry
 - (Geroch '71, Yang '03, ...)

Geroch decomposition I

- consider spacetime (\mathcal{M}, g_{MN}) with $\dim \mathcal{M} = D$
- consider existence of Killing vectors ξ^M
- consider ξ^M are either timelike or spacelike everywhere
- denote manifold S with $\dim S = D - 1$
 - collection of all trajectories of ξ^a
 - $v \in S$ is a curve in \mathcal{M} with v tangent to ξ^M everywhere
- consider tensor $T \in \mathcal{M}$ and $\tilde{T} \in S$
 - if
$$\xi_{M_1} T^{M_1 \dots M_p}_{N_1 \dots N_q} = \xi^{N_q} T^{M_1 \dots M_p}_{N_1 \dots N_q} = 0$$
$$\mathcal{L}_\xi T^{M_1 \dots M_p}_{N_1 \dots N_q} = 0$$

then there is a one-to-one correspondence between T and \tilde{T}
- the entire tensor field algebra on S is completely and uniquely mirrored by tensor field algebra on \mathcal{M}

Geroch decomposition II

- metric on (S, h_{MN}) : $h_{MN} = g_{MN} - \lambda^{-1} \xi_M \xi_N$, with $\lambda = \xi^M \xi_M$
- projection operator: $h^M_N = \delta^M_N - \lambda^{-1} \xi^M \xi_N$
- consider twist 2-form (exemplarily in $D = 5$):

$$\omega_{MN} = \epsilon_{MNJKL} \xi^J \nabla^K \xi^L$$

with $\omega_{MN} = \omega_{[MN]}$, $\omega_{MN} \xi^M = 0$

- one can show: $\mathcal{L}_\xi \lambda = 0$, $\mathcal{L}_\xi \omega_{MN} = 0$
 $\Rightarrow \lambda \in S$ and $\omega_{MN} \in S$

Geroch decomposition III

- with above relations derive

$$D_{[M}\omega_{NK]} \sim \epsilon_{MNKL_1L_2}\xi^{L_1(5)}R^{L_2}_{L_3}\xi^{L_3} = 0 \quad (\text{in vacuum})$$

\Rightarrow locally \exists a 1-form $A_M \in S$ with

$$\omega_{MN} = (dA)_{MN} = 2D_{[M}A_{N]}$$

- construct

$${}^{(4)}R_{MN} \sim (h_{MN}, (dA)_{MN}, (d\lambda)_M)$$

$$D^K(dA)_{KM} \sim (dA)_{KM}D^K\lambda$$

$$D^M D_M \lambda \sim D^M \lambda D_M \lambda - (dA)_{MN}(dA)^{MN}$$

\Rightarrow pure gravity in $D = 5$ is equivalent to

GR in $D = 4$ coupled to scalar field λ and vector field A_M

GREAT!

Are we done?

NOT QUITE!

Dimensional reduction I

Goal:

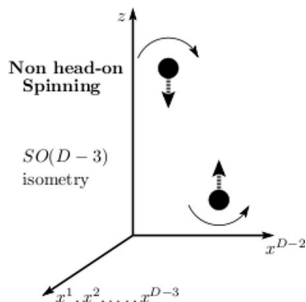
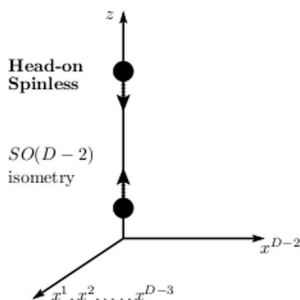
- study problems in $D \geq 5$
- reduction to $3 + 1$ problems

Why $3 + 1$?

- approaching limits of computational power
- we want numerical results in reasonable computational time
- we know how to do numerics in $3 + 1$
- vacuum GR in D dimensions
 \Rightarrow GR coupled to gauge and scalar fields in $D = 4$
- all quantities have geometrical interpretation
- straightforward modification of existing $3 + 1$ codes

Dimensional reduction II

- dimensional reduction by isometry on S^{D-4} sphere
 \Rightarrow has $SO(D-3) \subset SO(D-2)$ isometry group
- possible classes of models
 - in $D \geq 5$:
head-on collisions of non-spinning BHs
 - in $D \geq 6$:
BH collisions with impact parameter
collision of spinning BHs as long as dynamics are restricted to one plane



Zilhão et al '10

Dimensional reduction III

- dimensional reduction on S^{D-4} sphere
- ansatz for metric

Notation:

$$M, N, \dots = 0, \dots, D-1$$

$$\mu, \nu, \dots = 0, \dots, 3$$

$$\bar{\mu}, \bar{\nu}, \dots = 4, \dots, D-1$$

$$i, j, \dots = 1, 2, 3$$

$$\begin{aligned} ds^2 &= g_{MN} dx^M dx^N \\ &= g_{\mu\nu}(x^M) dx^\mu dx^\nu + \Omega_{\bar{\mu}\bar{\nu}}(x^M) (dx^{\bar{\mu}} - A^{\bar{\mu}}{}_\mu(x^M) dx^\mu) (dx^{\bar{\nu}} - A^{\bar{\nu}}{}_\nu(x^M) dx^\nu) \end{aligned}$$

Dimensional reduction IV

- $\{x^\mu\}$ coordinates in base space; $\{x^{\bar{\mu}}\}$ coordinates along the fibre
- consider Killing vector fields ξ^a with $a = 1, \dots, \frac{(D-4)(D-3)}{2}$

$$\mathcal{L}_\xi g_{MN} = 0$$

- Lie algebra $[\xi_a, \xi_b] = \epsilon_{ab}{}^c \xi_c$ with $\epsilon_{ab}{}^c$ structure constants of $SO(D-3)$
- fibre has minimal dimension necessary to host $\frac{(D-4)(D-3)}{2}$ independent KVs
- assume that KV has components exclusively along the fibre

$$\xi_a = \xi^{\bar{\mu}}{}_a \partial_{\bar{\mu}}$$

- normalize KVs, such that they only depend on coordinates along the fibre

$$\partial_\mu \xi^{\bar{\mu}}{}_a = 0$$

Dimensional reduction V

$\mathcal{L}_\xi g_{MN} = 0$ implies

① $M = \bar{\mu}, N = \bar{\nu}$: $\mathcal{L}_\xi g_{MN} = \mathcal{L}_\xi \Omega_{\bar{\mu}\bar{\nu}} = 0$

- $\Omega_{\bar{\mu}\bar{\nu}}$ is metric on maximally symmetric space $\rightarrow S^{D-4}$ sphere

$$\Omega_{\bar{\mu}\bar{\nu}} = \lambda(x^\mu) h_{\bar{\mu}\bar{\nu}}^{S^{D-4}}, \quad \text{with } h_{\bar{\mu}\bar{\nu}}^{S^{D-4}} \text{ metric on } S^{D-4} \text{ sphere}$$

② $M = \mu, N = \bar{\nu}$: $\mathcal{L}_\xi g_{MN} = -\mathcal{L}_\xi(\Omega_{\bar{\mu}\bar{\nu}} A^{\bar{\mu}}{}_\mu) = 0 \Rightarrow \mathcal{L}_\xi A^{\bar{\mu}}{}_\mu = 0$

- equivalent to $[\xi_a, A_\mu] = 0 \Rightarrow A^{\bar{\mu}}{}_\mu = 0$
- for $D \geq 5$ there are no non-trivial vector fields on S^{D-4} that commute with all KVs on the sphere
- all KVs are orthogonal to hypersurface

③ $M = \mu, N = \nu$: $\mathcal{L}_\xi g_{MN} = \mathcal{L}_\xi g_{\mu\nu} + \mathcal{L}_\xi(\Omega_{\bar{\mu}\bar{\nu}} A^{\bar{\mu}}{}_\mu A^{\bar{\nu}}{}_\nu) = 0 \Rightarrow \mathcal{L}_\xi g_{\mu\nu} = 0$

- KVs act transitively to the fibre
- base space metric $g_{\mu\nu}$ is independent of fibre coordinates

$$g_{\mu\nu}(x^M) = g_{\mu\nu}(x^\mu)$$

Dimensional reduction VI

- D dimensional spacetime metric has block diagonal form

$$ds^2 = g_{MN} dx^M dx^N = \underset{D=4}{g_{\mu\nu}} dx^\mu dx^\nu + \underset{D-4}{\lambda(x^\mu) h_{\bar{\mu}\bar{\nu}}^{S^{D-4}}} dx^{\bar{\mu}} dx^{\bar{\nu}}$$

- components of D dimensional Ricci tensor

① $M = \mu, N = \nu$:

$$^{(D)}R_{\mu\nu} = {}^{(4)}R_{\mu\nu} - \frac{D-4}{2\lambda} \left(\nabla_\mu \nabla_\nu \lambda - \frac{1}{2\lambda} \nabla_\mu \lambda \nabla_\nu \lambda \right)$$

② $M = \mu, N = \bar{\mu}$:

$$^{(D)}R_{\mu\bar{\mu}} = 0$$

③ $M = \bar{\mu}, N = \bar{\nu}$:

$$^{(D)}R_{\bar{\mu}\bar{\nu}} = {}^{(D-4)}R_{\bar{\mu}\bar{\nu}} - \frac{1}{2} h_{\bar{\mu}\bar{\nu}}^{S^{D-4}} \left(\nabla^\mu \nabla_\mu \lambda + \frac{D-6}{2\lambda} \nabla^\mu \lambda \nabla_\mu \lambda \right)$$

Dimensional reduction VII

- consider D dimensional vacuum spacetime $\Rightarrow {}^{(D)}R_{MN} = 0$
- equations of motion:

$$\mathcal{S} = \nabla^\mu \nabla_\mu \lambda + \frac{D-6}{2\lambda} \nabla^\mu \lambda \nabla_\mu \lambda - 2(D-5) = 0$$

$$E_{1,\mu\nu} = {}^{(4)}R_{\mu\nu} - \frac{1}{2}g_{\mu\nu} {}^{(4)}R - 8\pi T_{\mu\nu} = 0$$

with

$$T_{\mu\nu} = \frac{D-4}{16\pi\lambda} \left(\nabla_\mu \nabla_\nu \lambda - \frac{1}{2\lambda} \nabla_\mu \lambda \nabla_\nu \lambda - (D-5)g_{\mu\nu} + \frac{D-5}{4\lambda} g_{\mu\nu} \nabla^\kappa \lambda \nabla_\kappa \lambda \right)$$

- GR in D dimensions with $SO(D-3)$ isometry
 \Rightarrow reduction to GR in $D=4$ coupled (non-minimally) to scalar field

We are almost there!

3 + 1 decomposition I

- GR in D dimensions with $SO(D - 3)$ isometry
⇒ reduction to GR in $D = 4$ coupled (non-minimally) to scalar field
- reformulation as 3 + 1 time evolution problem
- 3 + 1 variables:
 - coordinates:
lapse function α
shift vector β^i
 - dynamical variables:
3-metric γ_{ij}
extrinsic curvature $K_{ij} = -\frac{1}{2\alpha}(\partial_t - \mathcal{L}_\beta)\gamma_{ij}$
scalar field λ
momentum of scalar field $K_\lambda = -\frac{1}{2\alpha}(\partial_t - \mathcal{L}_\beta)\lambda$
- 3 + 1 decomposition of GR coupled to scalar field
→ modify our mathematica notebook developed in the last lecture (◇)

Note: both notebooks will become available with the lecture notes and online

3 + 1 decomposition II

constraint equations

$$\mathcal{H} = R - K_{ij}K^{ij} + K^2$$

$$-\frac{D-4}{\lambda} \left(-(D-5) + D^i D_i \lambda + \frac{D-7}{4\lambda} D^i \lambda D_i \lambda - (D-5) \frac{K_\lambda^2}{\lambda} - 2KK_\lambda \right) = 0$$

$$\mathcal{M}_i = D_j K^j_i - D_i K - \frac{D-4}{2\lambda} \left(D_i K_\lambda - K^j_i D_j \lambda - \frac{K_\lambda}{\lambda} D_i \lambda \right) = 0$$

3 + 1 decomposition III

- solution to the initial data problem representing non-spinning BHs starting from rest
- time symmetric, i.e., modified Brill-Lindquist initial data
- assume maximal slicing ($K = 0$), conformal flatness

$$\gamma_{ij} = \psi^{\frac{4}{D-3}} \eta_{ij}, \quad K_{ij} = 0$$
$$\lambda = \psi^{\frac{4}{D-3}} y^2, \quad K_\lambda = 0$$

with

$$\psi_0 = 1 + \sum_{i=1}^N \frac{\mu_{(i)}}{4|r - r_{(i)}|^{D-3}}$$

- mass parameter

$$\mu_{(i)} \equiv r_{S_{(i)}}^{D-3} \equiv \frac{16\pi M_{(i)}}{\mathcal{A}_{D-2}(D-2)},$$

3 + 1 decomposition IV

evolution equations

$$(\partial_t - \mathcal{L}_\beta)\gamma_{ij} = -2\alpha K_{ij}$$

$$(\partial_t - \mathcal{L}_\beta)\lambda = -2\alpha K_\lambda$$

$$(\partial_t - \mathcal{L}_\beta)K_{ij} = -D_i D_j \alpha + \alpha (R_{ij} - 2K^k{}_i K_{kj} + K K_{ij})$$

$$- \alpha \frac{D-4}{2\lambda} \left(D_i D_j \lambda - \frac{1}{2\lambda} D_i \lambda D_j \lambda - 2K_\lambda K_{ij} \right)$$

$$(\partial_t - \mathcal{L}_\beta)K_\lambda = -\frac{1}{2} D^i \alpha D_i \lambda$$

$$+ \alpha \left((D-5) + K K_\lambda + \frac{D-6}{\lambda} K_\lambda^2 - \frac{1}{2} D^i D_i \lambda - \frac{D-6}{4\lambda} D^i \lambda D_i \lambda \right)$$

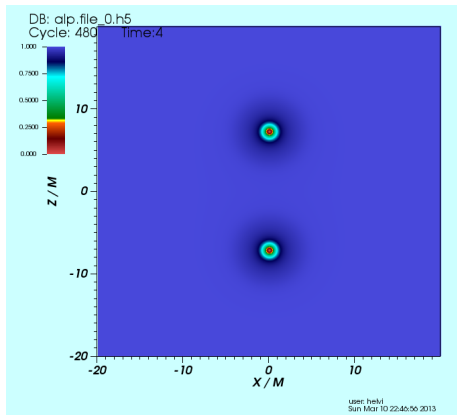
gauge choice – modified puncture gauge

$$\partial_t \alpha = \beta^i \partial_i \alpha - 2\alpha \left(K + (D-4)\mu_\lambda \frac{K_\lambda}{\lambda} \right)$$

$$\partial_t \beta^i = \beta^k \partial_k \beta^i - \eta_\beta \beta^i + \xi_\Gamma \tilde{\Gamma}^i + \xi_\lambda \frac{D-4}{2} \frac{\partial^i \lambda}{\lambda}$$

3 + 1 decomposition V

- choose strongly hyperbolic formulation, e.g., BSSN
(note: expressions become lengthy, see lecture notes)
- implementation / modification of 3 + 1 GR code
- example: **head-on collision** of two BHs starting from rest



Summary

- consider highly symmetric problems in GR in D dimensions
 \Rightarrow reduction to $3 + 1$ or $2 + 1$ system
- Cartoon method
 - use symmetry to reduce numerical grid
 - set up “slab” instead of “cube” \Rightarrow reduce from N^D grid points to $N^3 \cdot (2g + 1)$
 - use Cartesian coordinates \Rightarrow avoid coordinate singularities in grid
 - rotation and interpolation of functions from point $(*, X = 0, *)$ to $(*, X \neq 0, *)$
- dimensional reduction by isometry
 - consider problems with $SO(D - 3)$ isometry \Rightarrow reduction on S^{D-4} sphere
 - GR in D dimensions \Rightarrow GR in $D = 4$ coupled to scalar field
 - all quantities have geometric interpretation
 - straightforward modification of existing codes
- tomorrow: introduction to visualization toolkit VisIt